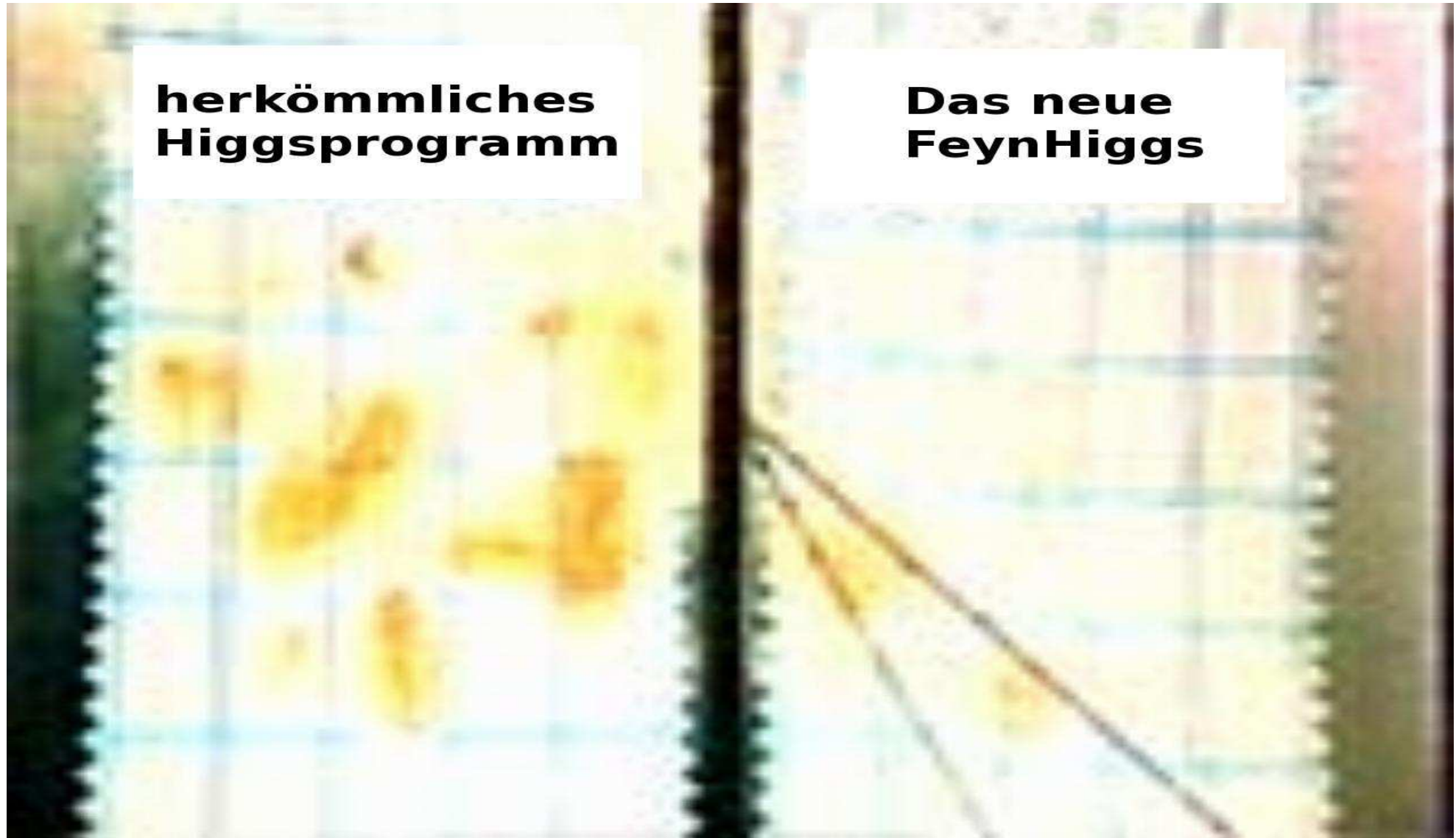


**herkömmliches  
Higgsprogramm**

**Das neue  
FeynHiggs**



# SUSY Prediction for Early LHC Data

*Sven Heinemeyer, IFCA (CSIC, Santander)*

CERN, 02/2009

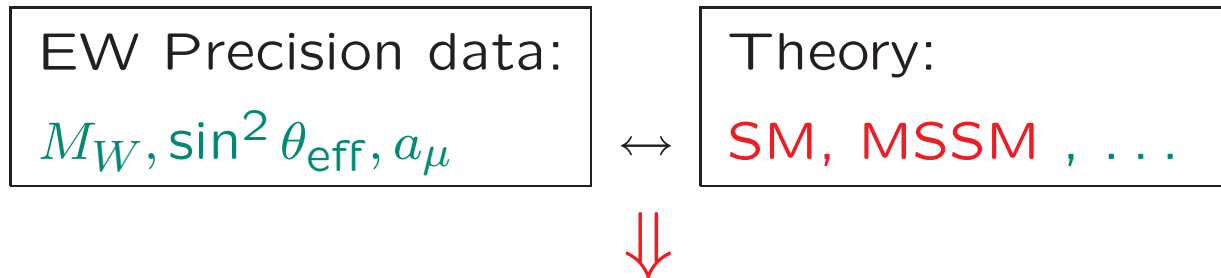
based on collaboration with

*O. Buchmüller, R. Cavanaugh, A. de Roeck, J. Ellis, G. Isidori,  
K. Olive, P. Paradisi, F. Ronga, G. Weiglein*

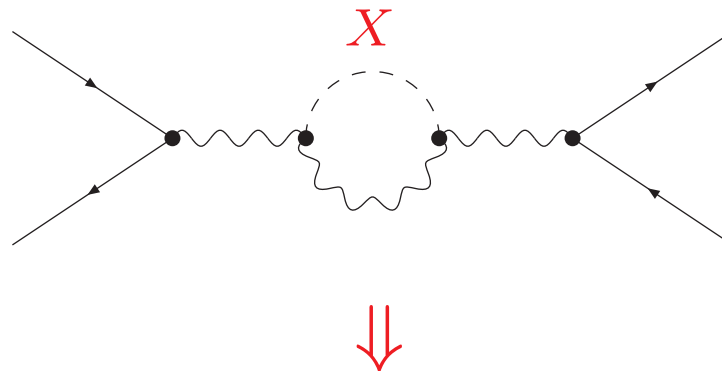
1. Introduction and motivation
2. Predictions for early LHC data
3. Conclusions

# 1. Introduction and motivation

Comparison of electro-weak precision observables with theory:



Test of theory at quantum level: Sensitivity to loop corrections

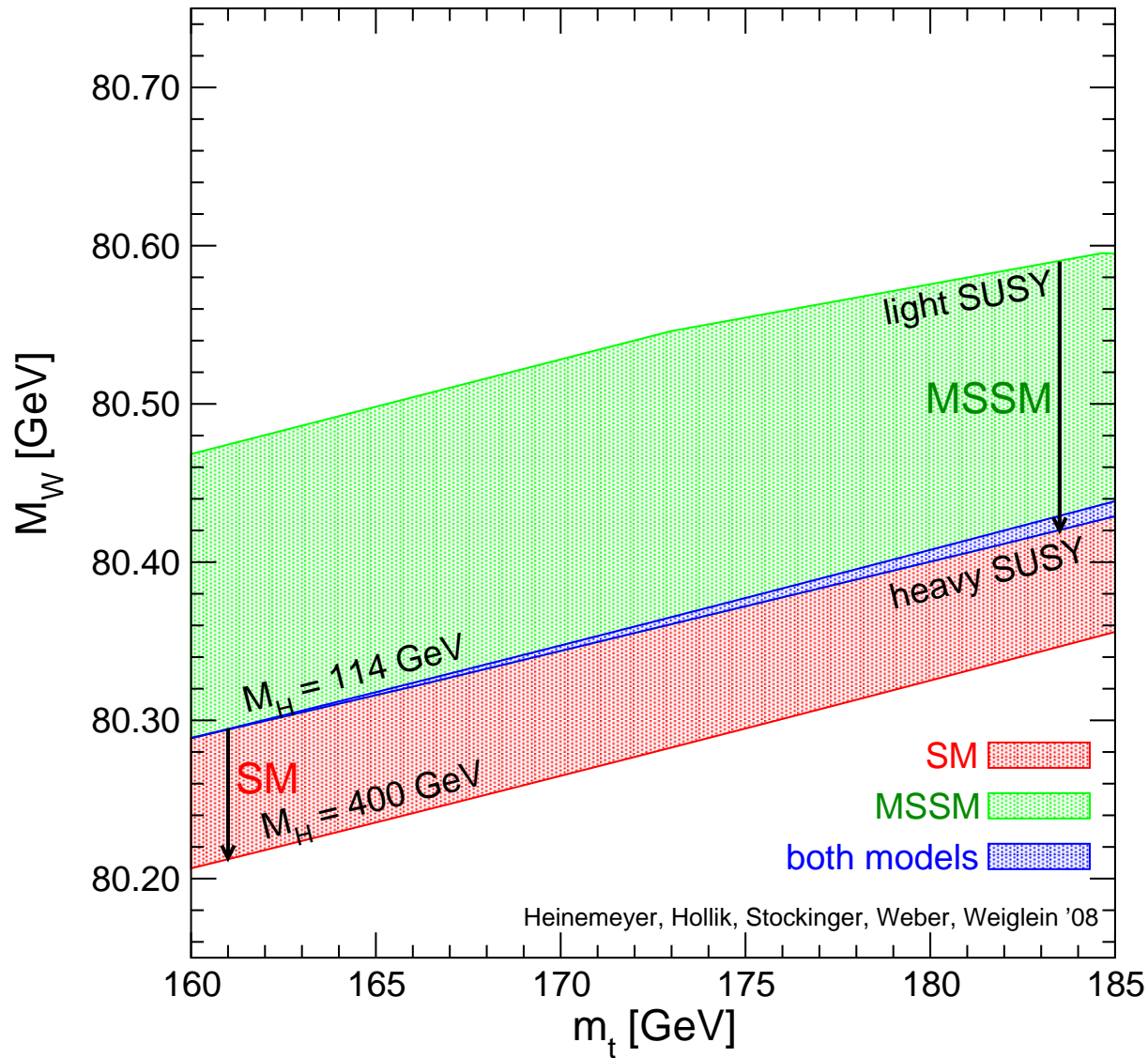


Very high accuracy of measurements and theoretical predictions needed

- Which model fits better?
- Does the prediction of a model contradict the experimental data?

Example: Prediction for  $M_W$  in the **SM** and the **MSSM** :

[S.H., W. Hollik, D. Stockinger, A.M. Weber, G. Weiglein '07]



**MSSM band:**

scan over  
SUSY masses

**overlap:**

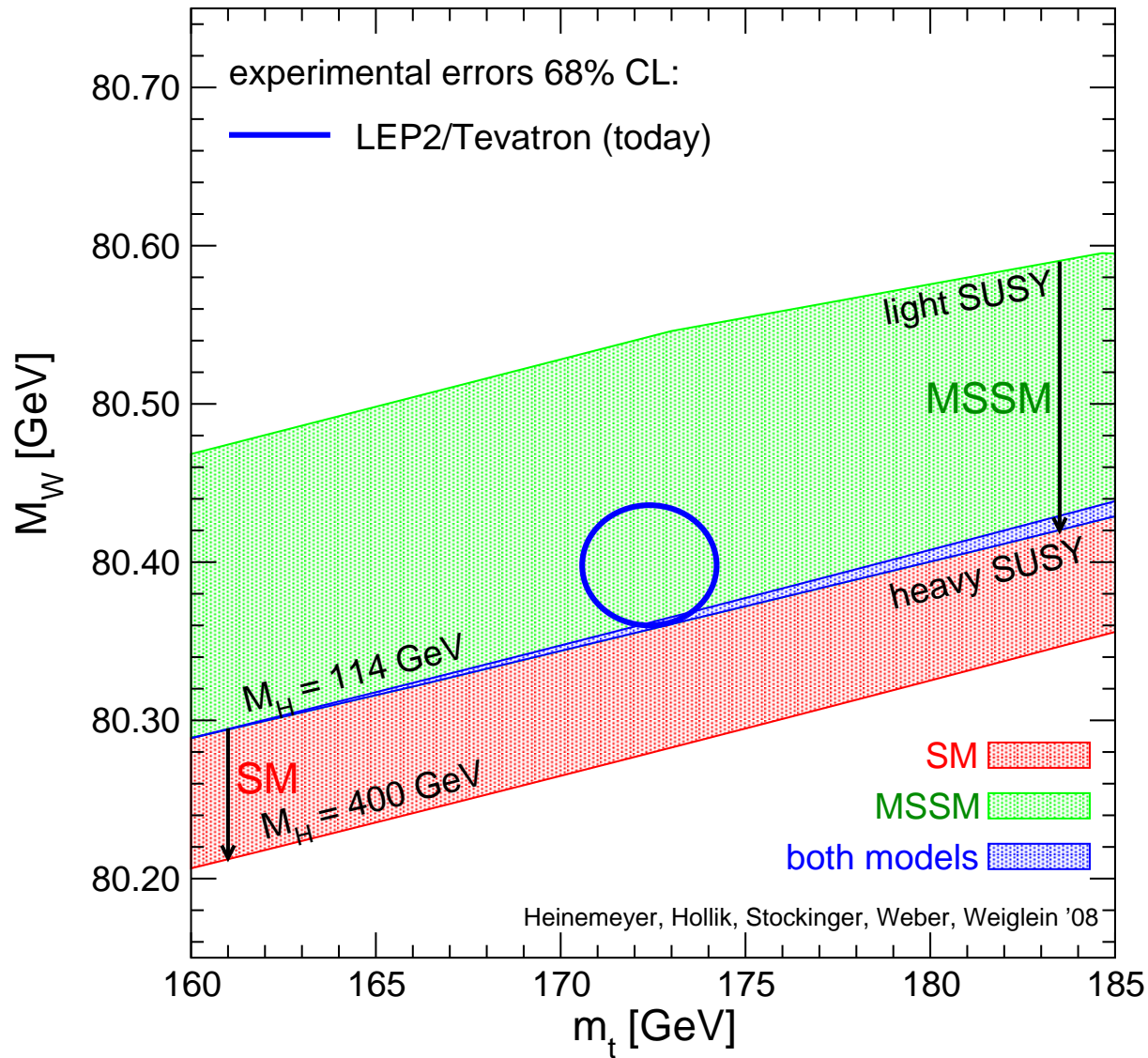
SM is MSSM-like  
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**SM band:**

variation of  $M_H^{\text{SM}}$

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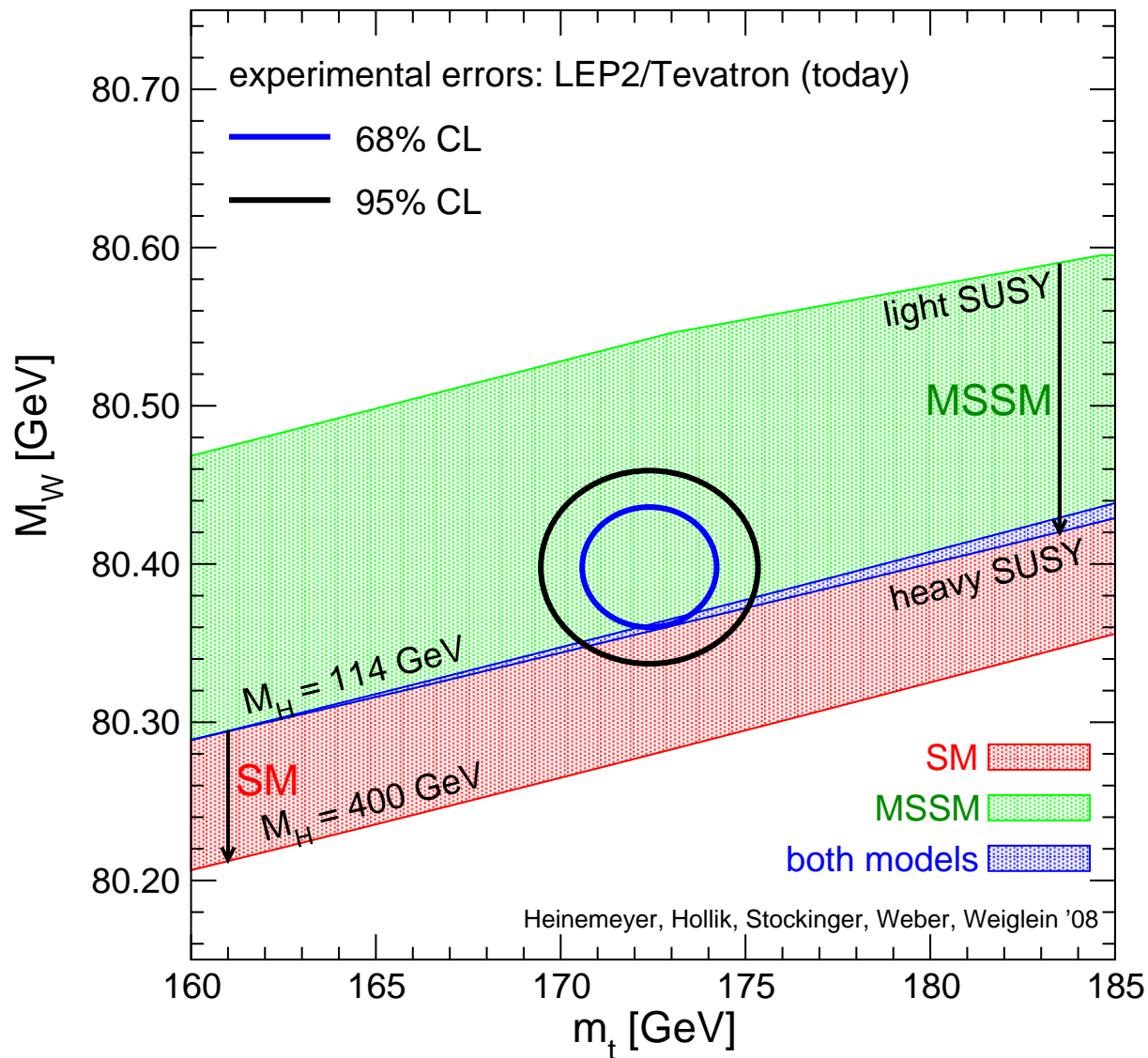
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## Indirect constraints on $M_{\text{SUSY}}$ from existing data?

- Electroweak precision observables (EWPO) ?
- $B$  physics observables (BPO) ?
- Cold dark matter (CDM) ?

⇒ combination of EWPO, BPO, CDM ?

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EWPO  $M_W$  : information on  $m_{\tilde{t}}$ ,  $m_{\tilde{b}}$  or  $M_A$ ,  $\tan \beta$  or ...

EWPO  $(g - 2)_\mu$  : information on  $\tan \beta$  and/or  $m_{\tilde{\chi}_0}$ ,  $m_{\tilde{\chi}^\pm}$  and/or  $m_{\tilde{\mu}}$ ,  $m_{\tilde{\nu}_\mu}$

BPO  $\text{BR}(b \rightarrow s\gamma)$  : information on  $\tan \beta$  and/or  $M_{H^\pm}$  and/or  $m_{\tilde{t}}$ ,  $m_{\tilde{\chi}^\pm}$

CDM (LSP gives CDM) : information on  $m_{\tilde{\chi}_1^0}$  and  $m_{\tilde{\tau}}$  or  $M_A$  or ...



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⇒ combination makes only sense if all parameters are connected!

⇒ GUT based models, ...

## Existing analyses for GUT based models: (involving precision observables)

### CMSSM/mSUGRA:

[J. Ellis, S.H., K. Olive, G. Weiglein '04, '06, '07] [J. Ellis, S.H., K. Olive, A.M. Weber, G. Weiglein '07]

[E. Baltz, P. Gondolo '04]

[R. de Austri, R. Trotta and L. Roszkowski '06, '07]

[B. Allanach, C. Lester and A.M. Weber '06, '07]

[O. Buchmueller et al. '07]

### NUHM (Non-Universal Higgs Mass model):

[J. Ellis, S.H., K. Olive, G. Weiglein '06]

[J. Ellis, S.H., K. Olive, A.M. Weber, G. Weiglein '07]

[J. Ellis, T. Hahn, S.H., K. Olive, G. Weiglein '07]

### VMSSM (Very Constrained MSSM):

[J. Ellis, S.H., K. Olive, G. Weiglein '06]

### mSUGRA (GDM) (Gravitino Dark Matter):

[J. Ellis, S.H., K. Olive, G. Weiglein '06]

### CMSSM, mGMSB, mAMSB:

[S.H., X. Miao, S. Su, G. Weiglein '08]

### Finite Unified Theories:

[S.H., M. Mondragón, G. Zoupanos '07]

## Different methods:

### 1.) Scanning:

- 3-dim scans (possibly with CDM fixing one dimension)

[J. Ellis, T. Hahn, SH, K. Olive, A.M. Weber, G. Weiglein '04, '06, '07]

- multi-dim scans

[O. Buchmueller et al. '07] [S.H., X. Miao, S. Su, G. Weiglein '08]

- multi-dim scans (with Markov Chain Monte Carlo technique)

[E. Baltz, P. Gondolo '04] [R. de Austri, R. Trotta and L. Roszkowski '06, '07]

[B. Allanach, C. Lester and A.M. Weber '06, '07] [O. Buchmueller et al. '08]

⇒ here: results using **last two**

### 2.) Fitting:

- Frequentist

[J. Ellis, T. Hahn, SH, K. Olive, A.M. Weber, G. Weiglein '04, '06, '07]

[O. Buchmueller et al. '07, '08] [S.H., X. Miao, S. Su, G. Weiglein '08]

- Bayesian (⇒ see talk by L. Roszkowski last week)

[R. de Austri, R. Trotta and L. Roszkowski '06, '07]

[B. Allanach, C. Lester and A.M. Weber '06, '07]

⇒ focus on **Frequentist** here

### 3.) Priors ... (flat)

## The models: 1.) CMSSM (or mSUGRA):

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$$

$m_0$  : universal scalar mass parameter

$m_{1/2}$  : universal gaugino mass parameter

$A_0$  : universal trilinear coupling

$\tan \beta$  : ratio of Higgs vacuum expectation values

$\text{sign}(\mu)$  : sign of supersymmetric Higgs parameter

} at the GUT scale

⇒ particle spectra from renormalization group running to weak scale

The models: 2.) NUHM1: (Non-universal Higgs mass model)

**Assumption:** no unification of scalar fermion and scalar Higgs parameter at the GUT scale

⇒ effectively  $M_A$  or  $\mu$  as free parameters at the EW scale

⇒ besides the CMSSM parameters

$M_A$  or  $\mu$

Further extension: NUHM2:

**Assumption:** no unification of the Higgs parameters at the GUT scale

⇒ effectively  $M_A$  and  $\mu$  as free parameters at the EW scale

⇒ besides the CMSSM parameters

$M_A$  and  $\mu$

## 2. Predictions for early LHC data

- combine all electroweak precision data as in the SM
- combine with  $B$  physics observables
- combine with CDM and  $(g - 2)_\mu$
- include SM parameters with their errors:  $m_t, M_Z, \Delta\alpha_{\text{had}}$

⇒  $\chi^2$  function

→ scan over the full CMSSM/NUHM1 parameter space

~  $2.5 \cdot 10^7$  points samples with MCMC

(comparison: L.R. et al.:  $0.04 \cdot 10^7$  points)

statistical measure:  $\chi^2$  function (Frequentist, no priors)

→ final minimum: Minuit

$\Delta\chi^2$ : 68, 95% C.L. contours

⇒ preferred CMSSM/NUHM1 parameters

⇒ LHC/ILC reach

## Best-fit points:

### CMSSM:

$$m_{1/2} = 310 \text{ GeV}, m_0 = 60 \text{ GeV}, A_0 = 240 \text{ GeV},$$

$$\tan \beta = 11, \mu = 380 \text{ GeV}, M_A = 410 \text{ GeV}$$

$$\chi^2/N_{\text{dof}} = 20.4/19 \text{ (37.3 \% probability)}$$

⇒ very similar to SPS 1a :-)

### NUHM1:

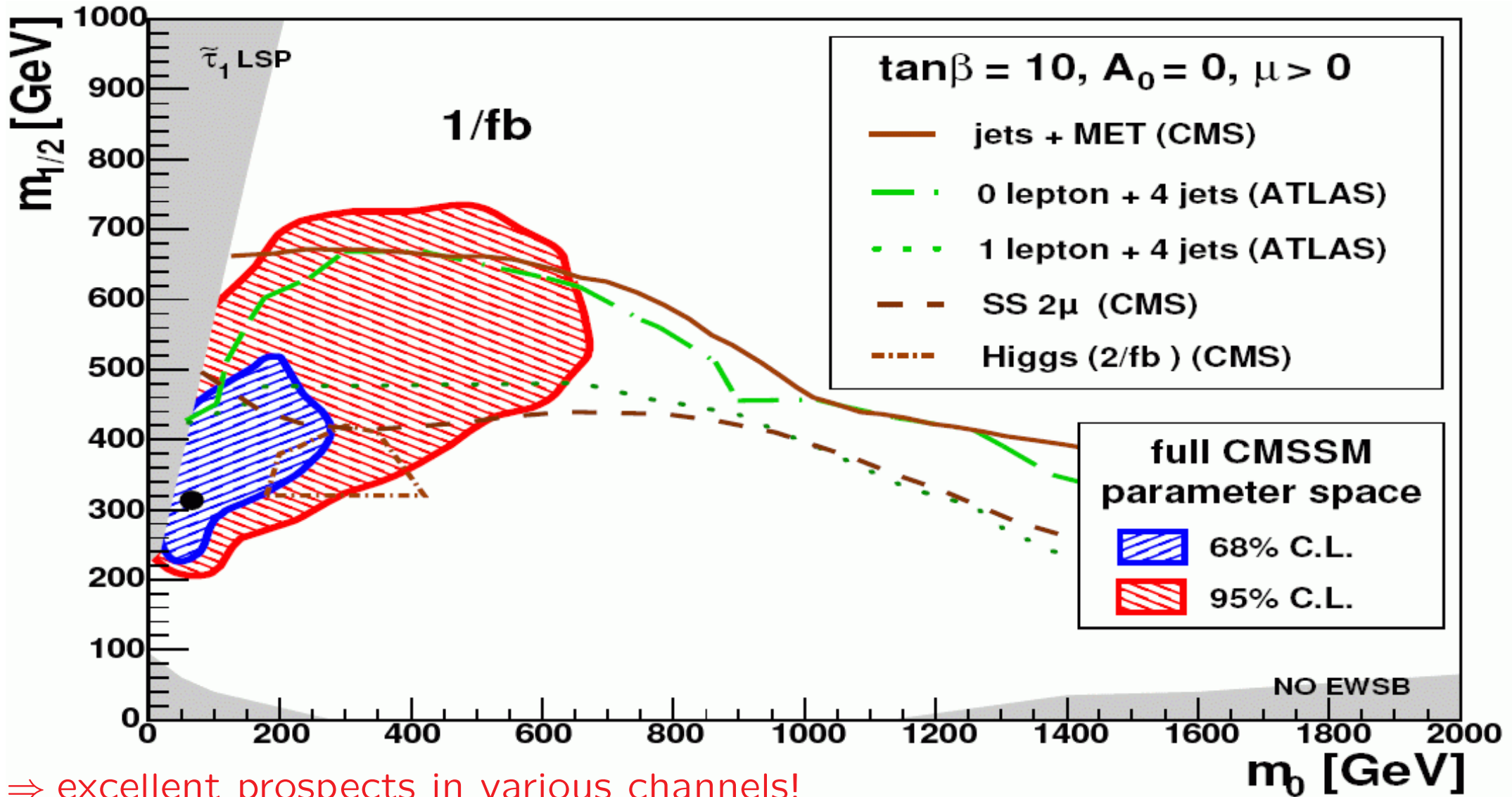
$$m_{1/2} = 240 \text{ GeV}, m_0 = 100 \text{ GeV}, A_0 = -930 \text{ GeV},$$

$$\tan \beta = 7, \mu = 870 \text{ GeV}, M_A = 300 \text{ GeV}$$

(39 % probability)

# LHC (CMS) reach with $1 \text{ fb}^{-1}$ :

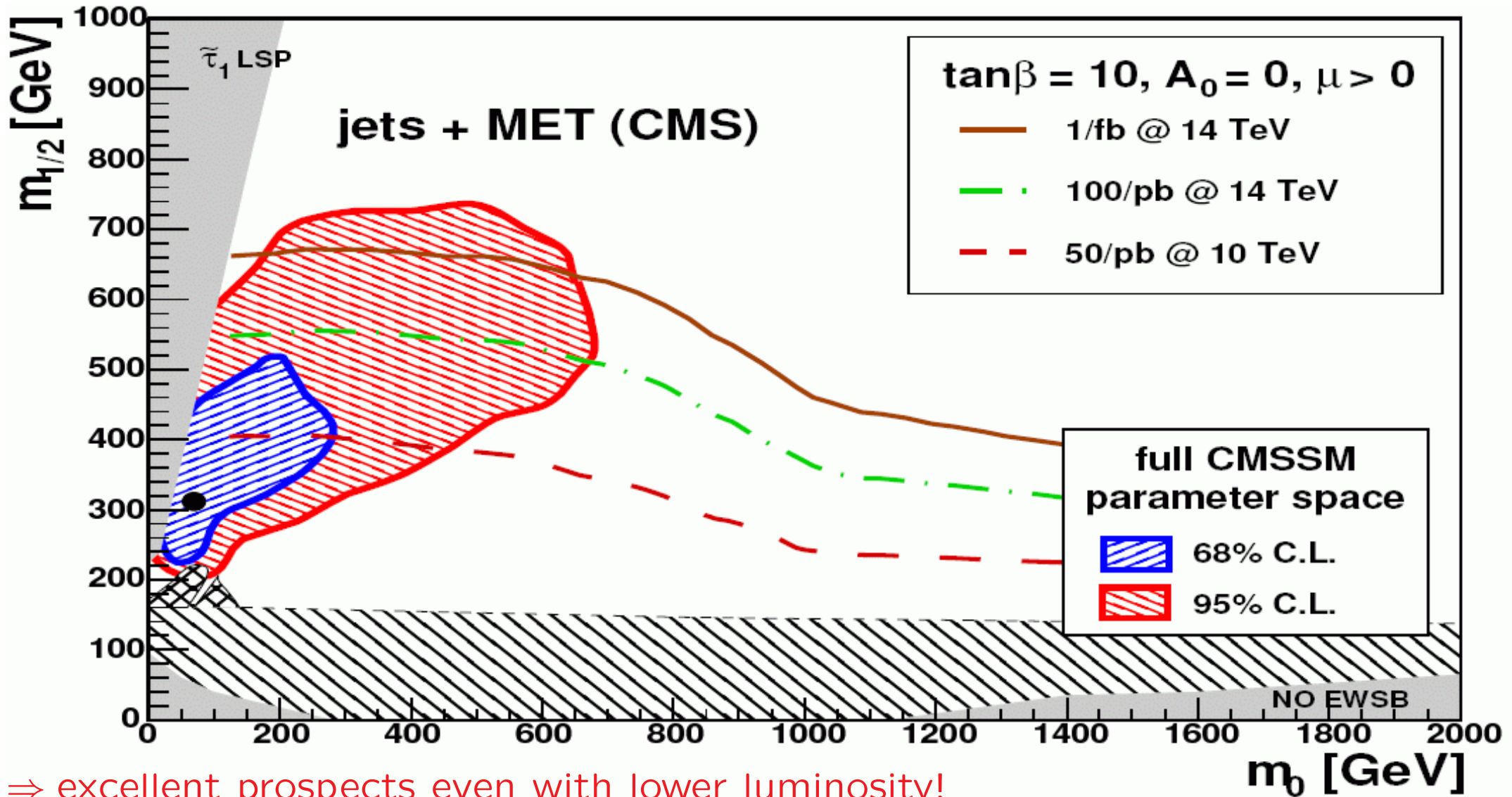
[CMS '07]





# LHC (CMS):

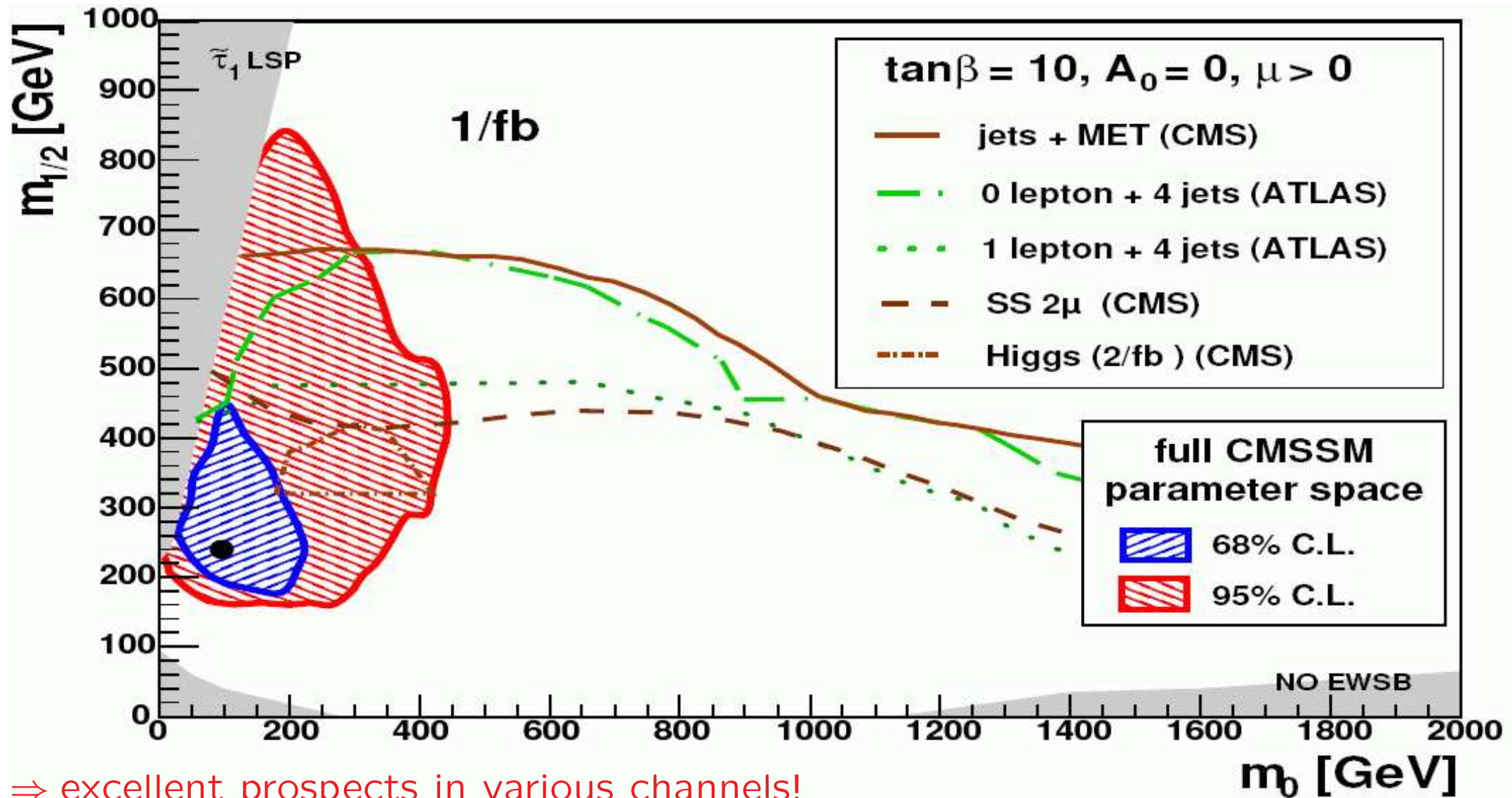
[CMS '07]



⇒ excellent prospects even with lower luminosity!

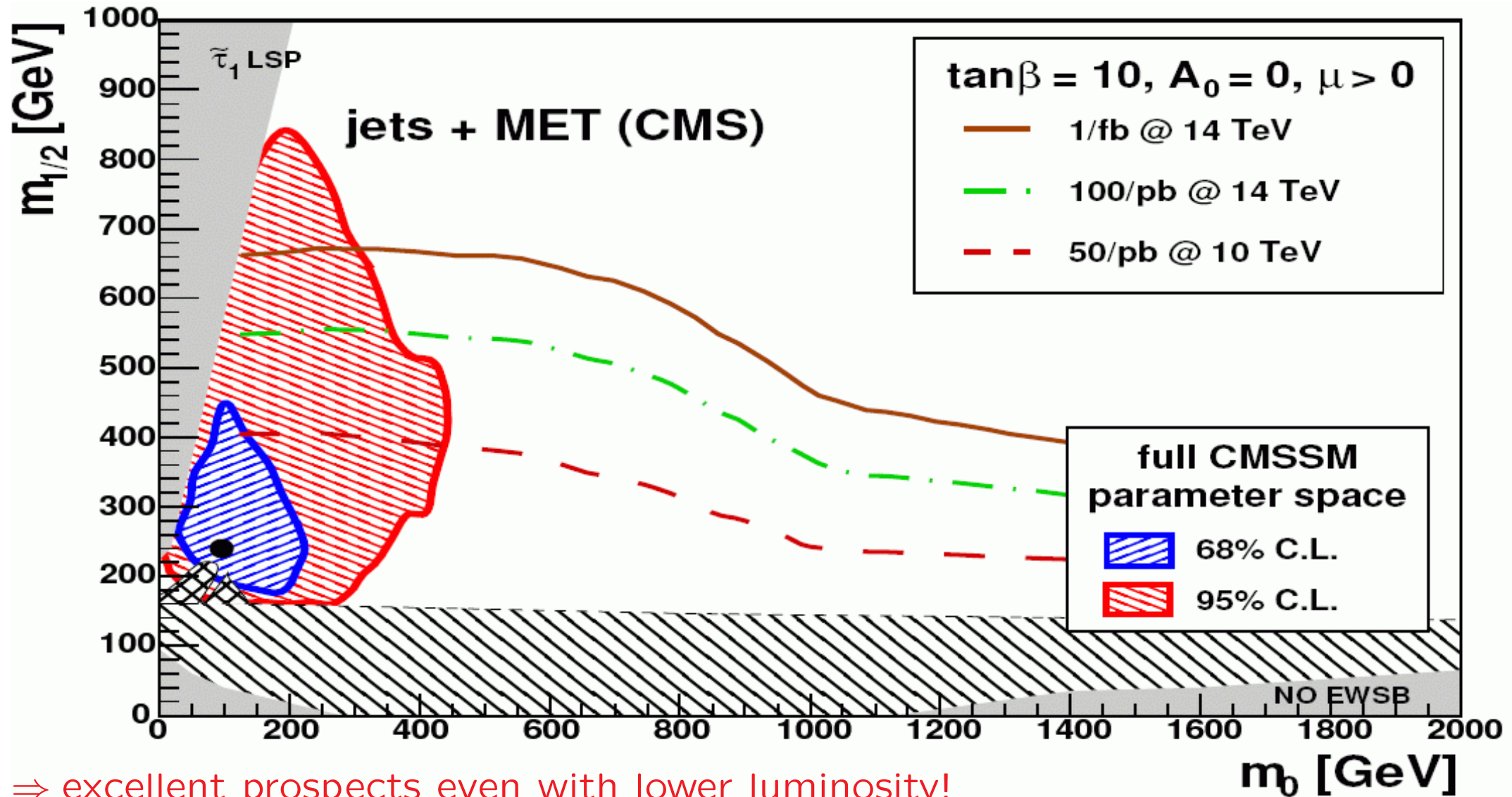
# LHC (CMS) reach with $1 \text{ fb}^{-1}$ : NUHM1 analysis

[CMS '07]



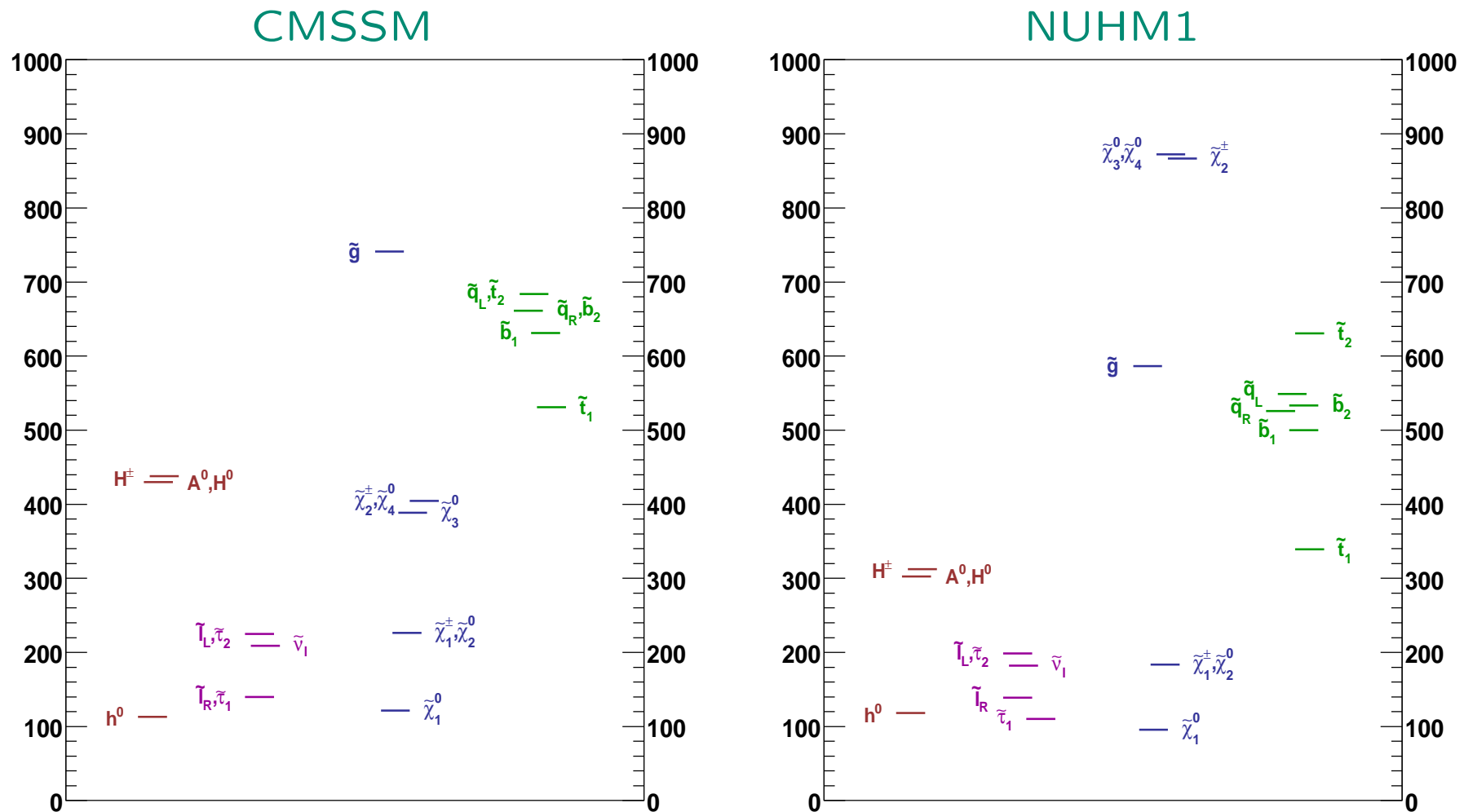
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[CMS '07]



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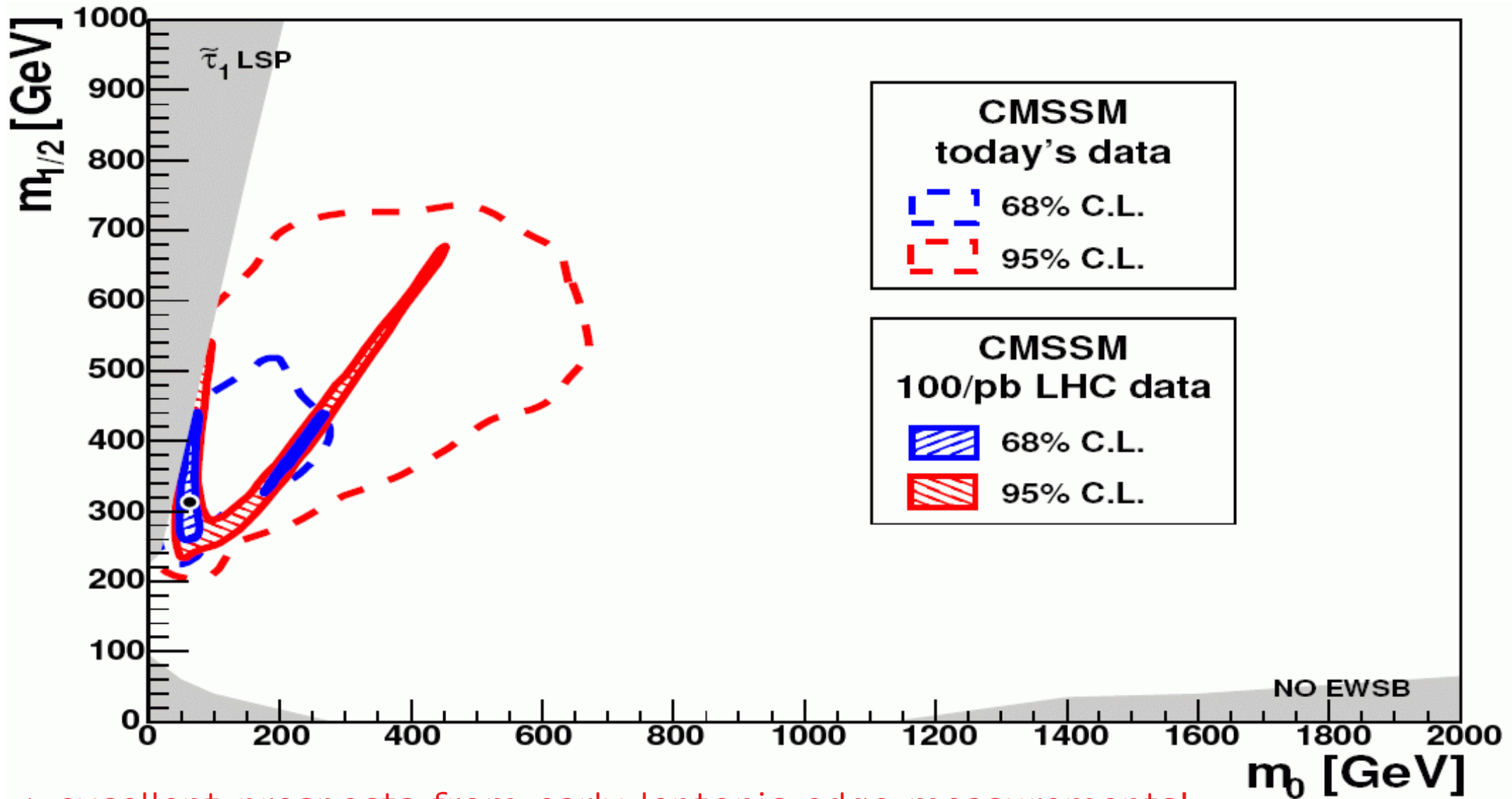
# Masses for best-fit points:



⇒ largely accessible spectrum for LHC and ILC

# LHC (CMS) reach with $1 \text{ fb}^{-1}$ :

[CMS '07] including leptonic edge measurements



⇒ excellent prospects from early leptonic edge measurements!



### 3. Conclusinos

- Idea: Predict most probable MSSM parameter regions using existing data: EWPO, BPO, CDM, ...
- Models: CMSSM, NUHM1
- statistical measure:  $\chi^2$  function (Frequentist, no priors)  
 $\sim 2.5 \cdot 10^7$  points samples with MCMC  
 $\Delta\chi^2$ : 68, 95% C.L. contours
- Best-fit points:  
CMSSM:  $m_{1/2} = 310$  GeV,  $m_0 = 60$  GeV,  $A_0 = 240$  GeV,  
 $\tan\beta = 11$ ,  $\mu = 380$  GeV,  $M_A = 410$  GeV  
 $\Rightarrow$  very similar to SPS 1a :-)  
NUHM1:  $m_{1/2} = 240$  GeV,  $m_0 = 100$  GeV,  $A_0 = -930$  GeV,  
 $\tan\beta = 7$ ,  $\mu = 870$  GeV,  $M_A = 300$  GeV
- 95% C.L. areas: mostly covered with  $\sim 1 \text{ fb}^{-1}$  (u.d.!)  
 $\Rightarrow$  early LHC data could be very conclusive!